## Position space renormalisation group for aggregation fractals

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## LETTER TO THE EDITOR

# Position space renormalisation group for aggregation fractals 

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#### Abstract

A simple position space renormalisation group scheme is used to find the fractal dimension of Sutherland type aggregation clusters grown in low-dimensional space. Clusters which move along paths of fractal dimension $d_{w}=0,1,2$ are considered and a systematic variation in cluster fractal dimension matching that of recent computer simulations is found.


Meakin (1983) and independently Kolb et al (1983) introduced a model of 'kinetic clustering of clusters' to describe the flocculation of colloids and aerosols. Thus far three different kinetics have been studied: clusters moving diffusively (Kolb et al 1983, Meakin 1983 and the experiments of Weitz and Oliveria 1984), clusters moving ballistically ( P Meakin, to be published) and the kinetics of chemically limited aggregation (CLA) (Schaeffer et al 1984, D A Weitz and M Oliveria, private communication). CLA may be realised as a computer simulation in which a pair of clusters is collided by positioning them repeatedly at random in a box until they are adjacent but not intersecting (W D Brown and R C Ball, to be published). These three kinetics correspond to requiring that clusters move along paths having fractal dimensions $d_{\mathrm{w}}=2,1,0$ respectively (Mandelbrot 1982).

A particularly simple cluster-cluster aggregation model is the hierarchical model due originally to Sutherland $(1967,1970)$ which has been studied numerically in low dimensions (Botet et al 1984, Jullien et al 1984a, b, W D Brown, private communication) and theoretically in high dimensions (Ball 1984, Ball and Witten 1984). In this model one starts with $N_{0}=2^{k}$ independent seed particles and constructs clusters iteratively. The ( $q+1$ th) iteration consists of forming $N_{0} / 2^{q+1}$ clusters each of $2^{q+1}$ particles. This is done by grouping the clusters formed in the $q$ th iteration into pairs and colliding each pair of clusters. This process continues until one cluster of $N_{0}$ particles is formed. We call a cluster of $2^{q}$ particles a $q$ cluster. All three kinetics, $d_{\mathrm{w}}=2,1,0$ have been considered for this model (Jullien et al 1984b, R Jullien and M Kolb, to be published, W D Brown, private communication). For ballistically moving clusters we consider only the case in which the impact parameter $b$ is chosen randomly. We do not consider the case in which two clusters are always collided along a line through their centres of mass, i.e. $b \equiv 0$ (Jullien 1984).

Aggregation clusters can be partly characterised by how the number of particles in the cluster, $N$, scales with the cluster radius of gyration, $R, N \sim R^{D(d)}$, in the limit $N \rightarrow \infty$. Here $D(d)$ is the cluster fractal dimension and $d$ is the dimension of Euclidean space in which the clusters form. Hentschel and Deutch (1984) suggested for Sutherland
clusters that

$$
\begin{array}{ll}
D(d)=d(4 d+3) /(9 d-2) & d<d_{\mathrm{c}}=8 \\
D(d)=d / 2 & d>d_{\mathrm{c}} \tag{2}
\end{array}
$$

independent of kinetics. This result was found using an argument which required two lengths, the cluster radius of gyration and the distance two clusters interpenetrate before a collision. It is presumably wrong: for low $d$ computer simulations show a systematic variation of $D(d)$ with kinetics (see table 1) and for high $d$ it has been shown (Ball and Witten 1984) that the exact answer is

$$
D(d)=D^{*}=\ln 4 / \ln (3 / 2) \quad \text { for } d>d_{\mathrm{c}}=2 D^{*}+d_{\mathrm{w}} .
$$

Table 1. Comparison of simulation results (Jullien et al $1984 \mathrm{~b}^{\mathrm{a}}$, R Jullien and $\mathrm{M} \mathrm{Kolb}{ }^{\mathrm{b}}$, to be published, WD Brown ${ }^{c}$, private communication) with position space renormalisation group (PSRG) results for fractal dimension $D(d)$ of Sutherland clusters moving with $d_{\mathrm{w}}=0,1,2$. PSRG results for $D\left(d_{\mathrm{c}}\right)$ where $d_{\mathrm{c}}=2 D^{*}+d_{\mathrm{w}}$ are given. These should be compared with the exact result $D\left(d_{\mathrm{c}}\right)=D^{*}$ for $d_{\mathrm{w}}=0,1,2$ where $D^{*}=\ln 4 / \ln (3 / 2)=3.42$ (Ball and Witten 1984).

| $d$ | $\begin{gathered} d_{\mathrm{w}}=0 \\ D(d) \end{gathered}$ |  | $\begin{gathered} d_{\mathrm{w}}=1 \\ D(d) \end{gathered}$ |  | $\begin{aligned} & d_{2}=2 \\ & D(d) \end{aligned}$ |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  | Simulation | PSRG | Simulation | PSRG | Simulation | PSRG |
| 1 |  | 1.00 |  | 0.89 |  |  |
| 2 | $1.53 \pm 0.04^{\text {b }}$ | 1.56 | $1.56 \pm 0.03^{\text {a }}$ | 1.41 | $1.42 \pm 0.03^{\text {a }}$ | 1.43 |
|  | $1.60 \pm 0.05^{\text {c }}$ |  |  |  |  |  |
| 3 | $1.98 \pm 0.04^{\text {b }}$ | 2.06 | $1.98 \pm 0.05^{\text {a }}$ | 1.87 | $1.78 \pm 0.05^{\text {a }}$ | 1.79 |
| 4 | $2.32 \pm 0.04^{\text {b }}$ | 2.53 | $2.35 \pm 0.08^{\text {a }}$ | 2.31 | $2.04 \pm 0.08^{\text {a }}$ | 2.13 |
| 5 |  | 3.03 | $2.5 \pm 0.2^{\text {a }}$ | 2.75 | $2.3 \pm 0.2^{\text {a }}$ | 2.46 |
| 6 |  | 3.44 | $2.7 \pm 0.2^{\text {a }}$ | 3.14 |  | 2.80 |
| 7 |  | 3.89 |  | 3.61 |  | 3.15 |
| 8 |  |  |  | 4.04 |  | 3.51 |
| 9 |  |  |  |  |  | 3.90 |
| $d_{\text {c }}$ |  | 3.82 |  | 3.97 |  | 3.84 |

To evaluate $D(d)$ for Sutherland clusters we use a simple position space renormalisation group scheme which requires only one length, the cluster radius of gyration. We first note that for Sutherland clusters $D(d)$ can be written (Botet et al 1984)

$$
\begin{aligned}
D(d) & =\lim _{q \rightarrow \infty} \frac{\ln \left(N_{q+1} / N_{q}\right)}{\ln \left(R_{q+1} / R_{q}\right)} \\
& =\lim _{q \rightarrow \infty} \frac{\ln 2}{\ln \lambda_{q}}
\end{aligned}
$$

where $R_{q}$ is the radius of gyration of a $q$ cluster, $N_{q}=2^{q}$ is the number of particles in a $q$ cluster and $\lambda_{q}=R_{q+1} / R_{q}$. The method we use for finding $\lambda_{q}$ is sufficiently simple to be independent of $q$. Thus we are forced to assume the method is valid for large $q$ and so gives a good approximation to $\lambda=\lim _{q \rightarrow \infty} \lambda_{q}$. We now write $\lambda$ in place of $\lambda_{q}$.

To find $\lambda$ we represent a $q$ cluster in $d$ dimensions by a sphere of radius $r_{q}$. This radius is determined by the requirement that the sphere must have the correct kinetic cross-section for hitting another $q$ cluster, itself represented by a sphere. In principle
we could model $r_{q}$ on any characteristic radius of the cluster which scales with the radius of gyration, but if a sphere is to be a good representation of a $q$ cluster we must choose the definition of $r_{q}$ which is most pertinent to the aggregation. We assume that for $d<d_{\mathrm{c}}$ there exists for a $q$ cluster a 'hitching radius' $R_{\mathrm{h}}(q)$ (Botet and Jullien 1984). This radius is defined to be the maximum radius that has the property that there is zero probability that the centres of mass of two collided $q$ clusters are less than $2 R_{\mathrm{h}}(q)$ apart. Thus $R_{\mathrm{h}}(q)$ defines an impenetrable hard sphere inside a $q$ cluster. As a low-dimensional approximation (exact in one dimension) we assume that the centres of two collided $q$ clusters are precisely $2 R_{\mathrm{h}}(q)$ apart. Of course actually the centres of mass of two collided $q$ clusters are at least $2 R_{\mathrm{n}}(q)$ apart from the definition of $R_{\mathrm{h}}(q) . R_{\mathrm{h}}(q)$ is the radius that we use for $r_{q}$. Using this approximation a $q+1$ cluster is represented as two touching spheres each of radius $r_{q}$. To find $r_{q+1}$, i.e. to find the correct radius for approximating (renormalising) two touching spheres of radius $r_{q}$ by a single sphere of radius $r_{q+1}$, we require that both the unrenormalised and renormalised representations of a $q+1$ cluster have the same cross-section to another $q+1$ cluster. The latter $q+1$ cluster is represented as a sphere of radius $r_{q+1}$ for simplicity. We write $\sigma_{1}\left(r_{q}, r_{+1}\right)$ for the cross-section of two touching spheres of radius $r_{q}$ to a single sphere of radius $r_{q+1}$ and we write $\sigma_{2}\left(r_{q}+1\right)$ for the cross-section of one sphere of radius $r_{q+1}$ to another sphere of radius $r_{q+1}$. Thus we require

$$
\sigma_{1}\left(r_{q}, r_{q+1}\right)=\sigma_{2}\left(r_{q+1}\right)
$$

or equivalently

$$
\begin{equation*}
\sigma_{1}(1, \lambda)=\sigma_{2}(\lambda) \tag{3}
\end{equation*}
$$

It is in the definitions of the cross-sections that cluster kinetics appear. Since solving (3) for $\lambda$ gives $D(d)$, the Sutherland cluster fractal dimension does depend on cluster kinetics in this approach in contrast to the earlier result of Hentschel and Deutch (1984) given by (1) and (2).

For the kinetics of cla excluded volume is used as the cross-section. That is, the cross-section of one cluster to another is taken to be the volume about one cluster in which the other cluster cannot sit without the two clusters intersecting. Thus $\sigma_{1}(1, \lambda)$ is the mutually excluded volume of two touching $d$-dimensional spheres of radius 1 (one $q+1$ cluster) and a sphere of radius $\lambda$ (the other, renormalised, $q+1$ cluster). This is the volume in $d$ dimensions of two intersecting spheres of radius $1+\lambda$ with centres 2 apart. Similarly $\sigma_{2}(\lambda)$ is the mutually excluded volume of two spheres of radius $\lambda$ which is the volume of a sphere of radius $2 \lambda$. With these definitions of $\sigma_{1}(1, \lambda)$ and $\sigma_{2}(\lambda)$ for the kinetics of CLA equation (3) reduces to (see figure 1)
$\left(\frac{2 \lambda}{1+\lambda}\right)^{d}=1+\left(\int_{0}^{\theta_{0}} \cos ^{d-1} \theta \mathrm{~d}(\sin \theta) / \int_{0}^{\pi / 2} \cos ^{d-1} \theta \mathrm{~d}(\sin \theta)\right)$
where

$$
\sin \theta_{0}=1 /(1+\lambda)
$$

The values obtained for $D(d)$ by solving (4) numerically are given in table 1 .
For Sutherland clusters which move linearly, the average ballistic cross-section is used for $\sigma_{1}$ and $\sigma_{2}$. Thus $\sigma_{1}(1, \lambda)$ is the average ballistic cross-section of two touching $d$-dimensional spheres of radius 1 to a sphere of radius $\lambda$. This is the isotropically averaged projected 'area' of two $d$-dimensional spheres of radius $1+\lambda$ with centres 2 apart. Similarly $\sigma_{2}(\lambda)$ is the average ballistic cross-section of one sphere of radius $\lambda$


Figure 1.


Figure 3.


Figure 2.

Figure 1. Geometry of excluded volume crosssection.

Figure 2. Geometry of average ballistic cross-section.
Figure 3. Geometry of diffusive cross-section.
to another sphere of radius $\lambda$ which is the isotropically averaged projected 'area' of a $d$-dimensional sphere of radius $2 \lambda$. This is of course the volume of a ( $d-1$ )dimensional sphere of radius $2 \lambda$. With these definitions of $\sigma_{1}(1, \lambda)$ and $\sigma_{2}(\lambda)$ for ballistically moving clusters equation (3) reduces to (see figure 2 )

$$
\begin{align*}
\left(\frac{2 \lambda}{1+\lambda}\right)^{d-1}= & 1+\left[\left(\int_{0}^{\pi / 2} \sin ^{d-2} \phi \int_{0}^{\theta_{0}(\phi)} \cos ^{d-2} \theta \mathrm{~d}(\sin \theta) \mathrm{d} \phi\right)\right. \\
& \left.\times\left(\int_{0}^{\pi / 2} \sin ^{d-2} \phi \int_{0}^{\pi / 2} \cos ^{d-2} \theta \mathrm{~d}(\sin \theta) \mathrm{d} \phi\right)^{-1}\right] \tag{5}
\end{align*}
$$

where

$$
\sin \theta_{0}(\phi)=(\sin \phi) /(1+\lambda) .
$$

The values obtained for $D(d)$ by solving (5) numerically are given in table 1 . The value $D(1)$ was obtained by solving (5) in the limit $d \rightarrow 1_{+}$.

For Sutherland clusters which move diffusively, $\sigma_{1}$ and $\sigma_{2}$ are diffusive crosssections. $\sigma_{1}$ is the steady state absorption rate by two touching spheres of radius 1 , of spheres of radius $\lambda$, from a field of unit concentration at infinity. That is,

$$
\sigma_{1}(1, \lambda)=\int_{S_{1}} \boldsymbol{\nabla} \phi_{1} \cdot \mathrm{~d} \boldsymbol{S} ; \quad \nabla^{2} \phi_{1}=0 ;\left.\quad \phi_{1}\right|_{s_{1}}=0 ; \quad \lim _{r \rightarrow \infty} \phi_{1}(r)=1 .
$$

Here $S_{1}$ is the surface of two intersecting spheres of radius $1+\lambda$ with centres 2 apart. Similarly $\sigma_{2}$ is the steady state absorption rate by a sphere of radius $\lambda$ of spheres of
radius $\lambda$ from a field of unit concentration at infinity. Thus we have

$$
\sigma_{2}(\lambda)=\int_{S_{2}} \nabla \phi_{2} \cdot \mathrm{~d} \boldsymbol{S} ; \quad \nabla^{2} \phi_{2}=0 ;\left.\quad \phi_{2}\right|_{S_{2}}=0 ; \quad \lim _{r \rightarrow \infty} \phi_{2}(\boldsymbol{r})=1
$$

where $S_{2}$ is the surface of a sphere of radius $2 \lambda$. To find $\lambda$ we use a crude approximation to $\phi_{1}$ to calculate $\sigma_{1}$. This is because we could not find an exact, tractable solution $\phi_{1}$ to Laplace's equation with the boundary conditions described above. The approximation used also has the advantage that it makes sense for non-integral dimensions. We set (see figure 3)

$$
\phi_{1}(\boldsymbol{r})=A /\left|\boldsymbol{r}+\boldsymbol{a}_{x}\right|^{d-2}+A /\left|\boldsymbol{r}-\boldsymbol{a}_{x}\right|^{d-2}+B /|\boldsymbol{r}|^{d-2}+1 .
$$

We also set

$$
\phi_{2}(\boldsymbol{r})=(2 A+B) /|\boldsymbol{r}|^{d-2}+1
$$

where

$$
\begin{equation*}
(2 A+B)=-1 /(2 \lambda)^{d-2} \tag{6}
\end{equation*}
$$

so the requirements that the total steady state flux through $S_{1}$ and $S_{2}$ be equal and that $\left.\phi_{2}\right|_{s_{2}}=0$ are automatically satisfied. We have also satisfied $\nabla^{2} \phi_{1}=0, \lim _{r \rightarrow \infty} \phi_{1}(\boldsymbol{r})=$ $1, \nabla^{2} \phi_{2}=0$ and $\lim _{r \rightarrow \infty} \phi_{2}(r)=1$. However, the requirement that $\phi_{1} \mid s_{1}=0$ can only be approximately met. We determine the four unknowns $A, B, a_{x}$ and $\lambda$ by requiring (6) be satisfied, the zero potential surface of $\phi_{1}$ pass through the extremities of $S_{1}$ (points C on figure 3) and that this surface has the same values of $\partial^{2} x / \partial y^{2}$ and $\partial^{4} x / \partial y^{4}$ at points C as $S_{1}$. This criterion is used as, for low $d$, most diffusing clusters hit the cluster tips. Thus we require at points $C$

$$
\begin{aligned}
-\frac{\partial^{2} \phi / \partial y^{2}}{\partial \phi / \partial x}= & -(1+\lambda)^{-1}, \\
& -\frac{\partial^{4} \phi / \partial y^{4}}{\partial \phi / \partial x}+\frac{6\left(\partial^{2} \phi / \partial y^{2}\right)\left(\partial^{3} \phi / \partial x \partial y^{2}\right)}{(\partial \phi / \partial x)^{2}}-\frac{3\left(\partial^{2} \phi / \partial y^{2}\right)^{2}\left(\partial^{2} \phi / \partial x^{2}\right)}{(\partial \phi / \partial x)^{3}} \\
= & -3(1+\lambda)^{-3},
\end{aligned}
$$

$\left.\phi_{1}\right|_{\mathrm{C}}=0$ and that (6) is satisfied. The values for $D(d)$ obtained by solving these equations numerically are given in table 1 . The value for $D(2)$ was obtained by solving them in the limit $d \rightarrow 2_{+}$.

We see from table 1 that for each $d$ the values for $D(d)$ obtained from simulations (Jullien et al 1984b, R Jullien and M Kolb, to be published, W D Brown, private communication) decrease as $d_{\mathrm{w}}$ increases as do the values for $D(d)$ obtained from the above position space renormalisation group approach. This systematic variation in the simulation values of $D(d)$ is expected since as $d_{\mathrm{w}}$ increases clusters interpenetrate less before colliding and thus form more open structures. We also see that for each of $d_{\mathrm{w}}=0,1,2$ the agreement between values for $D(d)$ obtained using this approach and those obtained from computer simulations worsens as $d$ increases. This is expected as we have used the low-dimensional approximation that $q$ clusters collide when their centres are precisely $2 r_{q}$ apart. The close agreement for low $d$ for $d_{\mathrm{w}}=2$ is no doubt a coincidence due to the rather ad hoc approximation used for $\phi_{1}$. However it should be noted that as $d \rightarrow \infty$ this approximation becomes an exact solution (with $B=0$ ). It can be shown analytically that as $d \rightarrow \infty, D(d)$ slowly approaches from above the
asymptote $\left(d-d_{\mathrm{w}}\right) / 2$ rather than the correct asymptote $D(d)=\ln 4 / \ln (3 / 2)$. This shows the method is self consistent; we made the low-dimensional approximation that Sutherland $q$ clusters interpenetrate only to $r_{q}$ and for all values of $d$ we find that $D(d)>\left(d-d_{\mathrm{w}}\right) / 2$ i.e. that Sutherland clusters are indeed opaque to each other for each of the three kinetics considered.

Work is continuing on the application of the above position space renormalisation group approach to polydispersed models of cluster-cluster aggregation (P Meakin, to be published) and to diffusion limited aggregation (Witten and Sander 1981).

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